

THE RELATIVISTIC UNIFORM MODEL: THE METRIC OF THE COVARIANT THEORY OF GRAVITATION INSIDE A BODY

S.G. Fedosin

Perm, Perm Territory, Russian Federation
sergey.fedosin@gmail.com

In the paper, it has been established that the sum of stress-energy tensors of the electromagnetic and gravitational fields, the acceleration and the pressure ones inside a stationary uniform spherical body vanishes within the framework of relativistic uniform model. This fact significantly simplifies a solution of the equation for the metric in the covariant theory of gravitation (CTG). The metric tensor components inside the body were calculated, and then they were combined with those of external metric tensor on the body's surface. The latter procedure also allowed us to exactly determine one of two unknown coefficients in the metric outside the body. The comparison between the CTG metric and the Reissner – Nordström one in general theory of relativity clearly demonstrated their difference caused by discrepancy between equations for the metric and a difference in formulations of the cosmological constant.

Keywords: metric, covariant theory of gravitation, scalar curvature, cosmological constant

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О МЕТРИКЕ КОВАРИАНТНОЙ ТЕОРИИ ГРАВИТАЦИИ ВНУТРИ ТЕЛА В РЕЛЯТИВИСТСКОЙ ОДНОРОДНОЙ МОДЕЛИ

С.Г. Федосин

г. Пермь, Пермский край, Российская Федерация
sergey.fedosin@gmail.com

В работе доказывається, що сума тензорів енергії-імпульса електромагнітного і гравітаційного полів, поля прискорень і поля тиску всередині нерухомого однорідного сферического тіла обертається в нуль в рамках релятивістської однорідної моделі. Це обставина суттєво спрощує рішення рівняння для метрики в коваріантній теорії гравітації (КТГ). Висліджуються компоненти метричного тензора всередині розглянутого тіла, а потім на його поверхні вони «сшиваються» з компонентами зовнішнього метричного тензора. Остання процедура дозволяє точно визначити один з двох невідомих коефіцієнтів в метриці за межами тіла. Порівняння метрики КТГ з метрикою Рейснера – Нордстрёма в загальній теорії відносності наочно показує їх різницю, яке обумовлено несовпадінням рівнянь для метрики, а також різницею формулювань космологічної постійної.

Ключевые слова: метрика, ковариантная теория гравитации, скалярная кривизна, космологическая постоянная, релятивистская система

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Introduction

In modern physics, the space-time metric of a certain physical system is completely defined by the corresponding metric tensor. The metric definition is of particular importance in the general theory of relativity, where the metric describes an action of gravity. In contrast, in the covariant theory of gravitation (CTG), gravity is an independent physical interaction. In this case, the metric of CTG is required mainly to describe the additional effects, associated with the interaction of electromagnetic waves with the gravitational field in the processes of space-time measurements by means of these waves. Accordingly, the metric form depends significantly on the theory of gravitation used.

Despite the success of the general theory of relativity in describing various gravitational phenomena, the theoretical foundation of this theory is still unsatisfactory. First of all, this is due to the absence of a generally recognized energy-momentum tensor of the gravitational field itself, the search for which has continued to this day [1 – 3]. Accordingly, the energy and momentum of a system becomes ambiguous or not conserved [4 – 6]. Other problems include emerging singularities, the need to interpret the cosmological constant, dark matter, dark energy, etc. In this regard, the search for alternatives to the general theory of relativity remains relevant, in particular, among vector-tensor theories [7 – 9].

The CTG refers to vector theories and has a well-defined energy-momentum tensor of the gravitational field. Outside a fixed spherical body, the metric tensor components within the framework of CTG were determined in Ref. [10]. Only the gravitational and electromagnetic fields exist outside the body, therefore only these fields exert their influence on the space-time metric here. Using this metric, it was possible to calculate the Pioneer effect, which has no explanation in the general theory of relativity [11]. CTG formulas describing the gravitational time dilation, the gravitational redshift of the wavelength, the signal delay in the gravitational field, lead to the same results as the general theory of relativity [12].

Next, we will calculate the metric of CTG inside a spherical body. In the presence of the matter, we should take into account the pressure field, which we consider in a covariant form as a vector field. Similarly, the concept of the vector acceleration field [13, 14] is used to calculate the energy and momentum of the matter, and its contribution into the equation for the metric. It is the representation of these fields in the form of vector fields that made it possible to find a covariant expression for the Navier – Stokes equation [15]. In contrast, in the general relativity, the pressure field and the acceleration one are almost always considered as simple scalar fields. Consequently, we can assume that CTG represents the contribution of the fields to the energy and momentum more accurately, as well as it does to the metric of the system.

In order to simplify the solution of the problem, we will assume that the matter of the body moves chaotically in the volume of the spherical shape, and it is kept from disruption by gravitation. The gravitational force in such macroscopic objects, as planets and stars, is so strong that it is sufficient to form the spherical shapes of them. This force is counteracted by the pressure force in the matter and the force from the acceleration field. One of the manifestations of the force from the acceleration field is the centrifugal force arising from that component of the particles' velocity, which is perpendicular to the radius-vector of the particles. We can also take into account the electromagnetic field and the corresponding force, which usually leads to repulsion of the charged matter in case of the excess charge of one sign. We will also assume that the physical system under consideration is a relativistic uniform system, in which the mass and charge distributions are similar to each other. This will allow us to use the expressions found earlier for the potentials and field strengths.

The need to determine the metric inside the matter arises as a consequence of the fact that the comparison of expressions for the components of the metric tensor inside and outside the matter makes it possible to unambiguously determine one of the unknown coefficients in the external metric. As a result, we obtain a more accurate expression for the CTG metric, suitable for solving more complex problems and considering small gravitational effects.

The equation for the metric

The use of the principle of least action leads to the following equation for the metric in CTG [14]:

$$R_{\alpha}^{\beta} - \frac{1}{4} R \delta_{\alpha}^{\beta} = -\frac{1}{2ck} (U_{\alpha}^{\beta} + W_{\alpha}^{\beta} + B_{\alpha}^{\beta} + P_{\alpha}^{\beta}), \quad (1)$$

where c is the speed of light; k is the constant, which is part of the Lagrangian in the terms with the scalar curvature R and with the cosmological constant Λ ; R_{α}^{β} is the Ricci tensor with the mixed indices; δ_{α}^{β} is the unit tensor (the Kronecker symbol); U_{α}^{β} , W_{α}^{β} , B_{α}^{β} and P_{α}^{β} are the stress-energy tensors of the gravitational, electromagnetic, the acceleration and the pressure ones, respectively.

As was shown in Ref. [16] all the quantities in Eq. (1) should be averaged over the volume of the system's typical particles, if Eq. (1) is used to find the metric inside the body. We will further assume that such averaging has already been carried out in Eq. (1). Another conclusion in Ref. [16] is that, within the framework of the relativistic uniform model, the scalar curvature inside a stationary body with the constant relativistically invariant mass density and charge is a certain constant quantity \bar{R} . In this case, the relation $\bar{R} = 2\bar{\Lambda}$ holds in CTG, where $\bar{\Lambda}$ is the averaged cosmological constant for the matter inside the body.

Acting in the same way as we did in Ref. [10], we will use the spherical coordinates

$$x^0 = ct, x^1 = r, x^2 = \theta, x^3 = \varphi,$$

related to the Cartesian coordinates by the relations:

$$x = r \sin \theta \cos \varphi, y = r \sin \theta \sin \varphi, z = r \cos \theta.$$

For the static metric, the standard form of the metric tensor of the spherical uniform body is as follows:

$$g_{\alpha k} = \begin{pmatrix} B & 0 & 0 & 0 \\ 0 & -K & 0 & 0 \\ 0 & 0 & -E & 0 \\ 0 & 0 & 0 & -E \sin^2 \theta \end{pmatrix}, \quad (2)$$

$$g^{k\beta} = \begin{pmatrix} \frac{1}{B} & 0 & 0 & 0 \\ 0 & -\frac{1}{K} & 0 & 0 \\ 0 & 0 & -\frac{1}{E} & 0 \\ 0 & 0 & 0 & -\frac{1}{E \sin^2 \theta} \end{pmatrix}, \quad (3)$$



where B, K, E are the functions of the radial coordinate r only and do not depend on the angular variables.

And there are four nonzero components of the metric tensor:

$$g_{00} = B, g_{11} = -K, g_{22} = -E, g_{33} = -E \sin^2 \theta.$$

By definition, the Christoffel coefficients $\Gamma_{\mu\nu}^{\beta}$ are expressed in terms of the metric tensor and its derivatives:

$$\Gamma_{\mu\nu}^{\beta} = \frac{1}{2} g^{\beta\gamma} (\partial_{\mu} g_{\gamma\nu} + \partial_{\nu} g_{\gamma\mu} - \partial_{\gamma} g_{\mu\nu}). \quad (4)$$

If we denote the derivatives with respect to the radius r by primes, then the nonzero Christoffel coefficients, expressed in terms of the functions B, K, E in the metric tensors (2) and (3), are equal, according to Eq. (4), to the following:

$$\begin{aligned} \Gamma_{01}^0 = \Gamma_{10}^0 = \frac{B'}{2B}, \quad \Gamma_{00}^1 = \frac{B'}{2K}, \quad \Gamma_{11}^1 = \frac{K'}{2K}, \quad \Gamma_{22}^1 = -\frac{E'}{2K}, \quad \Gamma_{33}^1 = -\frac{E' \sin^2 \theta}{2K}, \\ \Gamma_{12}^2 = \Gamma_{21}^2 = \Gamma_{13}^3 = \Gamma_{31}^3 = \frac{E'}{2E}, \quad \Gamma_{33}^2 = -\sin \theta \cos \theta, \quad \Gamma_{23}^3 = \Gamma_{32}^3 = \text{ctg} \theta. \end{aligned} \quad (5)$$

With the help of coefficients (5) we will calculate the components of the Ricci tensor with the covariant indices using the standard formula:

$$R_{\mu\nu} = \partial_{\alpha} \Gamma_{\mu\nu}^{\alpha} - \partial_{\nu} \Gamma_{\mu\alpha}^{\alpha} + \Gamma_{\mu\nu}^{\alpha} \Gamma_{\alpha\beta}^{\beta} - \Gamma_{\mu\beta}^{\alpha} \Gamma_{\alpha\nu}^{\beta}.$$

This will give four nonzero components:

$$\begin{aligned} R_{00} = \frac{B''}{2K} - \frac{B'^2}{4BK} - \frac{B'K'}{4K^2} + \frac{B'E'}{2KE}, \quad R_{11} = -\frac{B''}{2B} + \frac{B'^2}{4B^2} - \frac{E''}{E} + \frac{E'^2}{2E^2} + \frac{B'K'}{4BK} + \frac{K'E'}{2KE}, \\ R_{22} = -\frac{E''}{2K} + \frac{E'K'}{4K^2} - \frac{E'B'}{4BK} + 1, \quad R_{33} = \sin^2 \theta R_{22}. \end{aligned} \quad (6)$$

Eq. (1) contains the components of the Ricci tensor with the mixed indices, which can be found by multiplying the components of this tensor with the covariant indices by the metric tensor using the $R_{\alpha}^{\beta} = R_{\alpha\mu} g^{\mu\beta}$ formula. By application of components (6) and metric tensor (3), we find:

$$\begin{aligned} R_0^0 = \frac{B''}{2BK} - \frac{B'^2}{4B^2K} - \frac{B'K'}{4BK^2} + \frac{B'E'}{2BKE}, \\ R_1^1 = \frac{B''}{2BK} - \frac{B'^2}{4B^2K} - \frac{B'K'}{4BK^2} + \frac{E''}{KE} - \frac{E'^2}{2KE^2} - \frac{K'E'}{2K^2E}, \\ R_2^2 = \frac{E''}{2KE} + \frac{B'E'}{4BKE} - \frac{K'E'}{4K^2E} - \frac{1}{E}, \quad R_3^3 = R_2^2. \end{aligned} \quad (7)$$

Using formulas (6) and (3), we will calculate the scalar curvature as follows:

$$R = R_{\mu\nu} g^{\mu\nu} = \frac{B''}{BK} - \frac{B'^2}{2B^2K} - \frac{B'K'}{2BK^2} + \frac{B'E'}{BKE} + \frac{2E''}{KE} - \frac{E'^2}{2KE^2} - \frac{K'E'}{K^2E} - \frac{2}{E}. \quad (8)$$

The field tensors

The stress-energy tensors of the gravitational field [17, 18], the electromagnetic, the acceleration and the pressure ones [14], located on the right-hand side of the Eq. (1) for the metric, can be expressed as follows:

$$\begin{aligned}
 U_{\alpha}^{\beta} &= -\frac{c^2}{4\pi G} g^{\mu\kappa} \left(-\delta_{\alpha}^{\lambda} g^{\sigma\beta} + \frac{1}{4} \delta_{\alpha}^{\beta} g^{\sigma\lambda} \right) \Phi_{\mu\lambda} \Phi_{\kappa\sigma}, \\
 W_{\alpha}^{\beta} &= \varepsilon_0 c^2 g^{\mu\kappa} \left(-\delta_{\alpha}^{\lambda} g^{\sigma\beta} + \frac{1}{4} \delta_{\alpha}^{\beta} g^{\sigma\lambda} \right) F_{\mu\lambda} F_{\kappa\sigma}, \\
 B_{\alpha}^{\beta} &= \frac{c^2}{4\pi\eta} g^{\mu\kappa} \left(-\delta_{\alpha}^{\lambda} g^{\sigma\beta} + \frac{1}{4} \delta_{\alpha}^{\beta} g^{\sigma\lambda} \right) u_{\mu\lambda} u_{\kappa\sigma}, \\
 P_{\alpha}^{\beta} &= \frac{c^2}{4\pi\sigma} g^{\mu\kappa} \left(-\delta_{\alpha}^{\lambda} g^{\sigma\beta} + \frac{1}{4} \delta_{\alpha}^{\beta} g^{\sigma\lambda} \right) f_{\mu\lambda} f_{\kappa\sigma}.
 \end{aligned} \tag{9}$$

Here $\Phi_{\mu\lambda}$, $F_{\mu\lambda}$, $u_{\mu\lambda}$ and $f_{\mu\lambda}$ are the tensors of the gravitational, the electromagnetic, the acceleration and the pressure fields, respectively; G , ε_0 , η and σ are the gravitational, the electric, the acceleration and the pressure fields' constants, respectively.

The stress-energy tensors in Eqs. (9) were derived from the principle of the least action under the assumption that all the physical fields in the system under consideration were described as vector fields that had their own 4-potentials [13]. Due to the fact that the field tensors have the same form, it was possible to combine all the fields into a single general field [19, 20].

Let us express the 4-potentials of the fields in terms of the corresponding scalar and vector potentials of these fields:

$$\begin{aligned}
 D_{\lambda} &= \left(\frac{\Psi}{c}, -\mathbf{D} \right) \text{ for the gravitational field,} \\
 A_{\lambda} &= \left(\frac{\Phi}{c}, -\mathbf{A} \right) \text{ for the electromagnetic field,} \\
 U_{\lambda} &= \left(\frac{\mathfrak{g}}{c}, -\mathbf{U} \right) \text{ for the acceleration field,} \\
 \pi_{\lambda} &= \left(\frac{\mathfrak{p}}{c}, -\mathbf{\Pi} \right) \text{ for the pressure field.}
 \end{aligned}$$

The gravitational tensor is defined as the 4-curl of the 4-potential [17]. Similarly, the electromagnetic tensor, the acceleration tensor and the pressure field tensor [14] are calculated and have the following form:

$$\begin{aligned}
 \Phi_{\mu\lambda} &= \nabla_{\mu} D_{\lambda} - \nabla_{\lambda} D_{\mu} = \partial_{\mu} D_{\lambda} - \partial_{\lambda} D_{\mu}, \quad F_{\mu\lambda} = \nabla_{\mu} A_{\lambda} - \nabla_{\lambda} A_{\mu} = \partial_{\mu} A_{\lambda} - \partial_{\lambda} A_{\mu}, \\
 u_{\mu\lambda} &= \nabla_{\mu} U_{\lambda} - \nabla_{\lambda} U_{\mu} = \partial_{\mu} U_{\lambda} - \partial_{\lambda} U_{\mu}, \quad f_{\mu\lambda} = \nabla_{\mu} \pi_{\lambda} - \nabla_{\lambda} \pi_{\mu} = \partial_{\mu} \pi_{\lambda} - \partial_{\lambda} \pi_{\mu}.
 \end{aligned} \tag{10}$$



In the system under consideration, the vector potentials \mathbf{D} , \mathbf{A} , \mathbf{U} and $\mathbf{\Pi}$ of all the fields are close to zero because of the random motion of the matter's particles. This is due to the fact that the vector potentials of individual particles are directed along the particles' velocities, and therefore they change each time as a result of interactions.

The global vector potential of each field inside the body is calculated as the vector sum of the corresponding vector potentials of the particles. At each time point, most of the particles in the system have oppositely directed velocities and vector potentials, so that the vector sum of these potentials tends to zero on the average. The more particles are present in the system, the more exactly the equality to zero holds for the global vector potentials of the fields. We will not also take into account the proper vector potentials of individual particles. As was shown in Ref. [21], the energy of the particles' motion arises due to all these potentials, which is approximately equal to their kinetic energy. Thus the inaccuracy, arising from equating the vector potentials \mathbf{D} , \mathbf{A} , \mathbf{U} and $\mathbf{\Pi}$ to zero, does not exceed the inaccuracy in the case when only the rest energy is taken into account in the system's energy and the kinetic energy of the particles is neglected.

As for the scalar field potentials ψ , φ , ϑ and \wp , in the static case for a stationary spherical body, they must depend only on the current radius r and must not depend on either time or angular variables.

Assuming that $\mathbf{D} \approx 0$ and neglecting the contribution of the vector potential \mathbf{D} , in the spherical coordinates

$$x^0 = ct, \quad x^1 = r, \quad x^2 = \theta, \quad x^3 = \varphi$$

we find, from tensors (10) and (3), the nonzero components of the gravitational tensor:

$$\Phi_{01} = -\Phi_{10} = -\frac{1}{c} \frac{\partial \psi}{\partial r} - \frac{1}{c} \frac{\partial D_r}{\partial t} \approx -\frac{1}{c} \frac{\partial \psi}{\partial r}. \quad (11)$$

In Eq. (11) the quantity D_r in the spherical coordinates is the projection of the vector potential on the radial component of the 4-dimensional coordinate system. In this case, the quantity

$$\Gamma_r = c\Phi_{01} = -\frac{\partial \psi}{\partial r} - \frac{\partial D_r}{\partial t}$$

is the projection of the gravitational field strength on the radial component of the coordinate system.

The nonzero components of the electromagnetic field tensor, the acceleration one, and the pressure field tensor are obtained similarly to Eq. (11):

$$F_{01} = -F_{10} = -\frac{1}{c} \frac{\partial \varphi}{\partial r} - \frac{1}{c} \frac{\partial A_r}{\partial t} \approx -\frac{1}{c} \frac{\partial \varphi}{\partial r}, \quad u_{01} = -u_{10} = -\frac{1}{c} \frac{\partial \vartheta}{\partial r} - \frac{1}{c} \frac{\partial U_r}{\partial t} \approx -\frac{1}{c} \frac{\partial \vartheta}{\partial r},$$

$$f_{01} = -f_{10} = -\frac{1}{c} \frac{\partial \wp}{\partial r} - \frac{1}{c} \frac{\partial \Pi_r}{\partial t} \approx -\frac{1}{c} \frac{\partial \wp}{\partial r}. \quad (12)$$

In the Minkowski space-time, the special theory of relativity is valid, so that the potentials and the field strengths can be calculated exactly. For the case of the relativistic uniform model, the field strengths, which are part of the field tensors' components inside a spherical body, in the static case have the following form [22]:

$$\begin{aligned}
 c\Phi_{01} &= -\frac{Gc^2\gamma_c}{\eta r^2} \left[\frac{c}{\sqrt{4\pi\eta\rho_0}} \sin\left(\frac{r}{c}\sqrt{4\pi\eta\rho_0}\right) - r \cos\left(\frac{r}{c}\sqrt{4\pi\eta\rho_0}\right) \right] \approx -\frac{4\pi G\rho_0\gamma_c r}{3}, \\
 cF_{01} &= \frac{\rho_{0q}c^2\gamma_c}{4\pi\epsilon_0\rho_0\eta r^2} \left[\frac{c}{\sqrt{4\pi\eta\rho_0}} \sin\left(\frac{r}{c}\sqrt{4\pi\eta\rho_0}\right) - r \cos\left(\frac{r}{c}\sqrt{4\pi\eta\rho_0}\right) \right] \approx \frac{\rho_{0q}\gamma_c r}{3\epsilon_0}, \\
 cu_{01} &= \frac{c^2\gamma_c}{r^2} \left[\frac{c}{\sqrt{4\pi\eta\rho_0}} \sin\left(\frac{r}{c}\sqrt{4\pi\eta\rho_0}\right) - r \cos\left(\frac{r}{c}\sqrt{4\pi\eta\rho_0}\right) \right] \approx \frac{4\pi\eta\rho_0\gamma_c r}{3}, \\
 cf_{01} &= \frac{\sigma c^2\gamma_c}{\eta r^2} \left[\frac{c}{\sqrt{4\pi\eta\rho_0}} \sin\left(\frac{r}{c}\sqrt{4\pi\eta\rho_0}\right) - r \cos\left(\frac{r}{c}\sqrt{4\pi\eta\rho_0}\right) \right] \approx \frac{4\pi\sigma\rho_0\gamma_c r}{3}.
 \end{aligned} \tag{13}$$

In Eqs. (13) γ_c is the Lorentz factor of the typical particles that are moving at the center of the body; ρ_0 and ρ_{0q} denote the invariant mass and charge densities of the typical particles, respectively.

These mass and charge densities are obtained in the reference frames, which are comoving with the particles. It follows from Eqs. (10) – (13) that the field tensors inside the body are proportional to each other:

$$-\frac{\Phi_{\mu\lambda}}{G} = \frac{4\pi\epsilon_0\rho_0 F_{\mu\lambda}}{\rho_{0q}} = \frac{u_{\mu\lambda}}{\eta} = \frac{f_{\mu\lambda}}{\sigma}. \tag{14}$$

Let us sum up all the stress-energy tensors in formulas (9) and use Eq. (14):

$$\begin{aligned}
 U_\alpha^\beta + W_\alpha^\beta + B_\alpha^\beta + P_\alpha^\beta &= \\
 &= \frac{c^2}{4\pi\eta^2} \left(-G + \frac{\rho_{0q}^2}{4\pi\epsilon_0\rho_0^2} + \eta + \sigma \right) g^{\mu\kappa} \left(-\delta_\alpha^\lambda g^{\sigma\beta} + \frac{1}{4} \delta_\alpha^\beta g^{\sigma\lambda} \right) u_{\mu\lambda} u_{\kappa\sigma}.
 \end{aligned} \tag{15}$$

As was found in Ref. [23] from the equation of the particles' motion and in Ref. [24] from the generalized Poynting theorem, the following condition holds for the sum of the field coefficients inside the body:

$$-G + \frac{\rho_{0q}^2}{4\pi\epsilon_0\rho_0^2} + \eta + \sigma = 0. \tag{16}$$

Substituting condition (16) into Eq. (15) we find out that the sum of the stress-energy tensors inside the body, which is in equilibrium, becomes equal to zero:

$$U_\alpha^\beta + W_\alpha^\beta + B_\alpha^\beta + P_\alpha^\beta = 0. \tag{17}$$

Relation (17) was also derived in Ref. [24]. Will the result of Eq. (17) change if we consider the situation in the curved space-time? In the physical system in the form of a spherical body, the space-time metric is static and depends only on the radial coordinate. Since the vector field potentials are assumed to be zero, the tensor of each field contains only two nonzero components, which are equal in the absolute value. Taking into account the metric of the curved space-time leads to the fact that the tensors' components of each field in Eq. (13) must be multiplied by the same function Z that depends on the metric tensor components. Just as the metric tensor components, this function will depend



only on the radial coordinate. In the flat Minkowski space-time this function must be equal to unity, $Z = 1$, so that Eq. (13) is satisfied, which does not contain Z function.

Indeed, the equations for calculating the tensors of all the vector fields coincide with each other in their form, according to Refs. [13, 18, 25], and hence, the field tensors can differ from each other only by the constant coefficients at constant mass density ρ_0 and charge density ρ_{0q} . Therefore, if we multiply the tensor of each field, found in the Minkowski space-time, by the same function Z , in order to find this tensor in the curved space-time, relation (14) would not change, and an additional factor would appear on the right-hand side of Eq. (15). Since condition (16) always holds true, then, in the system under consideration, the sum of the stress-energy tensors in Eqs. (15) and (17) will also be zero in the curved space-time.

Calculation of the metric inside the body

Eq. (1) for the metric, in view of Eq. (17), is significantly simplified:

$$R_{\alpha}^{\beta} - \frac{1}{4} R \delta_{\alpha}^{\beta} = 0.$$

Substituting here (7) and (8), we get three equations:

$$\frac{B''}{2BK} - \frac{B'^2}{4B^2K} - \frac{B'K'}{4BK^2} + \frac{B'E'}{2BKE} - \frac{E''}{KE} + \frac{E'^2}{4KE^2} + \frac{K'E'}{2K^2E} + \frac{1}{E} = 0, \quad (18)$$

$$\frac{B''}{2BK} - \frac{B'^2}{4B^2K} - \frac{B'K'}{4BK^2} - \frac{B'E'}{2BKE} + \frac{E''}{KE} - \frac{3E'^2}{4KE^2} - \frac{K'E'}{2K^2E} + \frac{1}{E} = 0, \quad (19)$$

$$\frac{B''}{2BK} - \frac{B'^2}{4B^2K} - \frac{B'K'}{4BK^2} - \frac{E'^2}{4KE^2} + \frac{1}{E} = 0. \quad (20)$$

Substituting Eq. (20) in Eqs. (18) and (19) gives the same equation:

$$\frac{E''}{KE} - \frac{E'^2}{2KE^2} - \frac{B'E'}{2BKE} - \frac{K'E'}{2K^2E} = 0, \text{ or } \frac{E''}{E'} - \frac{E'}{2E} - \frac{B'}{2B} - \frac{K'}{2K} = 0. \quad (21)$$

If we subtract Eq. (18) from (19), we will get Eq. (21) again. The latter can be easily integrated, because each term represents the derivative of the natural logarithm of the corresponding function:

$$E' = C_1 \sqrt{BKE}, \quad (22)$$

where C_1 is a certain constant.

We will now use the condition obtained in Ref. [16], according to which the scalar curvature inside the body must be a constant value $\bar{R} = C_2$. With the help of the scalar curvature (8) we obtain an expression:

$$\frac{B''}{BK} - \frac{B'^2}{2B^2K} - \frac{B'K'}{2BK^2} + \frac{B'E'}{BKE} + \frac{2E''}{KE} - \frac{E'^2}{2KE^2} - \frac{K'E'}{K^2E} - \frac{2}{E} = C_2. \quad (23)$$

The sum of Eqs. (23) and (18) gives the following one:

$$\frac{B''}{BK} - \frac{B'^2}{2B^2K} - \frac{B'K'}{2BK^2} + \frac{B'E'}{BKE} = \frac{C_2}{2}.$$

Comparing this expression with Eq. (20), we obtain:

$$\frac{E'^2}{4KE^2} - \frac{1}{E} + \frac{B'E'}{2BKE} = \frac{C_2}{4}.$$

Next, we will substitute here the value of $K = \frac{E'^2}{C_1^2 BE}$ according to Eq. (22):

$$\frac{E'}{E} = \frac{2C_1^2 B'}{4 + C_2 E - C_1^2 B}. \quad (24)$$

Now we need K from Eq. (22) and the relation $\frac{K'}{K}$ from Eq. (21):

$$K = \frac{E'^2}{C_1^2 BE}, \quad \frac{K'}{K} = \frac{2E''}{E'} - \frac{B'}{B} - \frac{E'}{E}. \quad (25)$$

Let us substitute expressions (25) into Eq. (20):

$$B'' - \frac{B'E''}{E'} + \frac{B'E'}{2E} - \frac{BE'^2}{2E^2} + \frac{2E'^2}{C_1^2 E^2} = 0. \quad (26)$$

Eqs. (24) and (26) together form a system of two differential equations in the functions B and E . Direct substitution shows us that the system of these equations has the following solution:

$$E = r^2, \quad B = \frac{C_3}{r} + \frac{4}{C_1^2} + \frac{C_2 r^2}{3C_1^2}.$$

Indeed, in the weak gravitational field, when the curved space-time turns into the Minkowski one, it should be $E = r^2$, $B = K \approx 1$ in the spherical coordinates.

In order to ensure that the function B is not infinitely large at the center at $r = 0$, the constant C_3 must be equal to zero. From the condition $B \approx 1$ it follows that $C_1 = 2$, and from Eq. (22) we get the equality $BK = 1$. In addition, the constant C_2 must be sufficiently small. As a result, for the metric tensor components we can write the following:

$$\begin{aligned} B = g_{00} &= 1 + \frac{C_2 r^2}{12}, \quad E = -g_{22} = r^2, \\ K = -g_{11} &= \frac{1}{B} = \frac{1}{1 + \frac{C_2 r^2}{12}}, \quad g_{33} = -r^2 \sin^2 \theta. \end{aligned} \quad (27)$$

The constant C_2 in expressions (27) represents the value of the scalar curvature, averaged over the volume of a typical particle, which is constant inside the body, so that $\bar{R} = C_2$.

In Ref. [16] we found the relation for the value of the cosmological constant $\bar{\Lambda}$ averaged over the volume of a typical particle:

$$-ck\bar{\Lambda} = \frac{G\rho_0}{a} \left[\frac{c^3\gamma_c}{\eta\sqrt{4\pi\eta\rho_0}} \sin\left(\frac{a}{c}\sqrt{4\pi\eta\rho_0}\right) - m_g \right] - \frac{\rho_{0q}}{4\pi\epsilon_0 a} \left[\frac{\rho_{0q}c^3\gamma_c}{\eta\rho_0\sqrt{4\pi\eta\rho_0}} \sin\left(\frac{a}{c}\sqrt{4\pi\eta\rho_0}\right) - q_b \right] + \rho_0\wp_c - \frac{\sigma\rho_0c^2\gamma_c}{\eta}.$$

Expanding the sine by the rule

$$\sin x \approx x - \frac{x^3}{6},$$

in view of Eq. (16), we find:

$$-ck\bar{\Lambda} \approx -\frac{Gm_g\rho_0}{a} - \frac{Gm\rho_0\gamma_c}{2a} + \rho_0c^2\gamma_c + \frac{q_b\rho_{0q}}{4\pi\epsilon_0 a} + \frac{q\rho_{0q}\gamma_c}{8\pi\epsilon_0 a} + \rho_0\wp_c, \quad (28)$$

$$C_2 = \bar{R} = 2\bar{\Lambda} \approx -\frac{2}{ck} \left(\rho_0\Psi_a - \frac{Gm\rho_0\gamma_c}{2a} + \rho_0c^2\gamma_c + \rho_{0q}\varphi_a + \frac{q\rho_{0q}\gamma_c}{8\pi\epsilon_0 a} + \rho_0\wp_c \right),$$

where k is the factor,

$$k = -\frac{c^3}{16\pi G\beta} \quad (29)$$

(β is a certain constant of the order of unity);

$\Psi_a = -\frac{Gm_g}{a}$ is the scalar potential of the gravitational field on the surface of the body at $r = a$;

$\varphi_a = \frac{q_b}{4\pi\epsilon_0 a}$ is the scalar potential of the electric field (a is the radius of the body); m_g , q_b are the

gravitational mass and the total charge of the body; γ_c is the Lorentz factor of the particles at the center of the body; \wp_c is the potential of the pressure field at the center of the sphere; the mass

$m = \frac{4\pi a^3\rho_0}{3}$ and the charge $q = \frac{4\pi a^3\rho_{0q}}{3}$ are auxiliary quantities.

In the brackets, on the right-hand side of Eq. (28), there is the sum of the volumetric energy densities of the particles in the scalar field potentials: the first and second terms are from the gravitational field, the third one is from the acceleration field, the fourth and fifth terms are from the electric field, and the sixth one is from the pressure field.

The third term is the greatest, it is proportional to the rest energy density of the body. If we take into account only this term, then, in the first approximation, the constant C_2 will be equal to

$$C_2 \approx \frac{32\pi G\rho_0\gamma_c\beta}{c^2}.$$

Comparison of the metric tensor components inside and outside the body

At $r = a$ the current radius reaches the surface of the spherical body, and here the internal metric becomes equal to the external one. It means that we can equate the components of the corresponding metric tensors at $r = a$. According to Ref. [10], the metric tensor components outside the body in the covariant theory of gravitation are equal to

$$g_{00} = 1 + \frac{A_3}{r} + \frac{1}{16\pi ckr^2} \left(Gm_g^2 - \frac{q_b^2}{4\pi\epsilon_0} \right), \quad g_{22} = -r^2, \quad (30)$$

$$g_{11} = -\frac{1}{1 + \frac{A_3}{r} + \frac{1}{16\pi ckr^2} \left(Gm_g^2 - \frac{q_b^2}{4\pi\epsilon_0} \right)}, \quad g_{33} = -r^2 \sin^2 \theta.$$

Comparison of Eqs. (30) and (27) shows that the components g_{22} and g_{33} coincide both inside and outside the body.

Equating g_{00} in Eqs. (27) and (30) under condition that $r = a$, taking into account Eqs. (28) and (29), we find the constant A_3 :

$$A_3 \approx \frac{2G\beta}{c^4} \left[mc^2 \gamma_c + \left(m - \frac{1}{2} m_g \right) \psi_a - \frac{Gm^2 \gamma_c}{2a} + \left(q - \frac{1}{2} q_b \right) \varphi_a + \frac{q^2 \gamma_c}{8\pi\epsilon_0 a} + m\phi_c \right]. \quad (31)$$

According to Ref. [22], the gravitational mass m_g of the body and the total electric charge q_b are determined as follows:

$$m_g = \int \rho_0 \gamma' dV =$$

$$= \frac{c^2 \gamma_c}{\eta} \left[\frac{c}{\sqrt{4\pi\eta\rho_0}} \sin\left(\frac{a}{c} \sqrt{4\pi\eta\rho_0}\right) - a \cos\left(\frac{a}{c} \sqrt{4\pi\eta\rho_0}\right) \right] \approx m\gamma_c \left(1 - \frac{3\eta m}{10ac^2} \right). \quad (32)$$

$$q_b = \int \rho_{0q} \gamma' dV =$$

$$= \frac{\rho_{0q} c^2 \gamma_c}{\eta\rho_0} \left[\frac{c}{\sqrt{4\pi\eta\rho_0}} \sin\left(\frac{a}{c} \sqrt{4\pi\eta\rho_0}\right) - a \cos\left(\frac{a}{c} \sqrt{4\pi\eta\rho_0}\right) \right] \approx q\gamma_c \left(1 - \frac{3\eta m}{10ac^2} \right).$$

Since $\gamma_c > 0$, it turns out that $m_g > m$ and $q_b > q$.

Now we will substitute Eq. (31) into the expression for g_{00} (see Eq. (30)) and take into account Eq. (29):

$$g_{00} = -\frac{1}{g_{11}} \approx 1 + \frac{2Gm\gamma_c\beta}{c^2 r} +$$

$$+ \frac{2G\beta}{c^4 r} \left[m\psi_a + \frac{1}{2} m_g (\psi - \psi_a) - \frac{Gm^2 \gamma_c}{2a} + q\varphi_a + \frac{1}{2} q_b (\varphi - \varphi_a) + \frac{q^2 \gamma_c}{8\pi\epsilon_0 a} + m\phi_c \right]. \quad (33)$$

In this expression $\psi = -\frac{Gm_g}{r}$ and $\varphi = \frac{q_b}{4\pi\epsilon_0 r}$ denote the scalar potentials of the gravitational and electric fields outside the body, respectively.



We can also determine the quantities γ_c and \wp_c more exactly. In Ref. [21] we found the expression for the square of the particles' velocities v_c^2 at the center of the spherical body; using it we can estimate the value of the Lorentz factor in Eq. (33):

$$\gamma_c = \frac{1}{\sqrt{1-v_c^2/c^2}} \approx 1 + \frac{v_c^2}{2c^2} + \frac{3v_c^4}{8c^4} \approx 1 + \frac{3\eta m}{10ac^2} \left(1 + \frac{9}{2\sqrt{14}}\right) + \frac{27\eta^2 m^2}{200a^2 c^4} \left(1 + \frac{9}{2\sqrt{14}}\right)^2.$$

According to Ref. [26], the scalar potential of the pressure field at the center of the body is approximately equal to

$$\wp_c \approx \frac{3\sigma m}{10a} \left(1 + \frac{9}{2\sqrt{14}}\right),$$

while the acceleration field constant η and the pressure field constant σ are given by the formulas

$$\eta = \frac{3}{5} \left(G - \frac{\rho_{0q}^2}{4\pi\epsilon_0\rho_0^2} \right), \quad \sigma = \frac{2}{5} \left(G - \frac{\rho_{0q}^2}{4\pi\epsilon_0\rho_0^2} \right).$$

In Eq. (33) we see the complex structure of the metric tensor components, in which additional terms appear as compared to the Minkowski spacetime metric, where in the spherical coordinates

$$g_{00} = -\frac{1}{g_{11}} = 1.$$

The main addition in Eq. (33) is the term

$$\frac{2Gm\gamma_c\beta}{c^2 r},$$

and if we take into account Eq. (32), then this addition will become approximately equal to $-\frac{2\psi\beta}{c^2}$.

The second important addition includes square brackets in Eq. (33), which, by the order of magnitude, determines the energy of the gravitational and electric fields, as well as the pressure one. In these brackets, we can also use the approximate relation of the masses in expression (32). For the metric tensor components outside the body all this leads to the following expression:

$$g_{00} = -\frac{1}{g_{11}} \approx 1 - \frac{2\psi\beta}{c^2} + \frac{2G\beta}{c^4 r} \left(m\psi_a + \frac{1}{2}m_g\psi + \frac{3\eta m^2\gamma_c}{10a} + q\varphi_a + \frac{1}{2}q_b\varphi + m\wp_c \right). \quad (34)$$

On the right-hand side of Eq. (34), in the round brackets, there are quantities with the dimension of energy. For large cosmic bodies, the main quantity here is the negative energy associated with gravitation. In this case we can see that the third term, containing c^4 in the denominator, is distinguished by a sign from the second term, containing c^2 in the denominator.

Comparison with the metric of the general theory of relativity

In order to compare with the metric tensor components (30) and (34), we will consider the Reissner – Nordström metric in the spherical coordinates, which describes the static gravitational field around a charged spherical body in the general theory of relativity. We will use our notation for the field potentials:

$$g_{00} = 1 + \frac{2\Psi}{c^2} + \frac{Gq_b\varphi}{c^4 r}, \quad g_{11} = -\frac{1}{1 + \frac{2\Psi}{c^2} + \frac{Gq_b\varphi}{c^4 r}}, \quad g_{22} = -r^2, \quad g_{33} = -r^2 \sin^2 \theta. \quad (35)$$

As we can see, the second and third terms in the component g_{00} in the Reissner – Nordström metric (35) differ significantly from the corresponding terms in the component g_{00} in the CTG metric (34) outside the body. For example, we can see that the metric in Eq. (35) does not reflect the energy of the pressure field inside the body in any way, whereas in Eq. (34) the energy $m\varphi_c$ is associated with the pressure field and makes its contribution to the metric. Taking into account Eq. (28), the energy $m\varphi_c$ also defines the metric (27) inside the body.

This difference in the form of the metric is due to the difference in the equations for determining the metric in both theories. While Eq. (1) is used in CTG, the equation for the metric with the cosmological constant Λ in the general theory of relativity has the following form in the matter with the stress-energy tensor T_α^β :

$$R_\alpha^\beta - \frac{1}{2} R \delta_\alpha^\beta + \Lambda \delta_\alpha^\beta = \frac{8\pi G}{c^4} T_\alpha^\beta. \quad (36)$$

According to the approach of the general theory of relativity, the action of gravitation must be described by the metric tensor, and therefore T_α^β does not include the stress-energy tensor of the gravitational field. There is no matter and no pressure field outside the charged body; only the electromagnetic field is left on the right-hand side of Eq. (36), so that we have $T_\alpha^\beta = W_\alpha^\beta$. As a rule, the term with the cosmological constant Λ in Eq. (36) is neglected due to its smallness, and then the solution for the metric (35) is obtained.

Since the cosmological constant is taken into account in CTG fully, it turns out that the solution of Eq. (27) in view of (28) for the CTG metric inside the body and the solution of Eq. (33) outside the body are more precise and informative than the solution of Eq. (35) in the Reissner – Nordström metric. Moreover, in CTG, the cosmological constant $\bar{\Lambda}$ is not equal to zero and is proportional to the potentials of all the fields acting inside the body. If in Eq. (28) only the main term with rest energy density is taken into account, then with the relation (29) we can estimate the value $\bar{\Lambda}$:

$$\bar{\Lambda} \approx -\frac{\rho_0 c \gamma_c}{k} \approx \frac{16\pi G \rho_0 \beta \gamma_c}{c^2}. \quad (37)$$

If we substitute here the average mass density of the cosmic space matter of the observable universe, we shall obtain the value $\bar{\Lambda} \approx 10^{-52} \text{ m}^{-2}$. The smallness of the cosmological constant $\bar{\Lambda}$ inside cosmic bodies is associated with the large factor (29) in Eq. (37). To this end we recall that the issue of the cosmological constant in the general theory of relativity has not yet been resolved unambiguously [27], especially with respect to correlation with vacuum energy. Here it is implied that a very large vacuum energy makes little contribution to the metric for some reason and to the small cosmological constant.

In CTG, the greater is the mass density in Eq. (37), the larger is Λ inside the body. However if we distribute the matter of all cosmic bodies over the space, then the mass density will be very low, which leads to insignificantly small value $\bar{\Lambda} \approx 10^{-52} \text{ m}^{-2}$. We should also pay attention to the fact that the cosmological constant outside the body is assumed to be zero due to its gauging in CTG [16]. Inside the bodies, as well as inside the observable universe as some global body, $\bar{\Lambda}$ has a certain value. In the approximation of the relativistic uniform body model, $\bar{\Lambda}$ is determined in Eq. (28).

In contrast, in the general theory of relativity, in Eq. (36), the nonzero value of the cosmological constant outside the body is admitted. The latter follows from the possibility of influence of the zero vacuum's energy on the metric through the cosmological constant.



Summary

In Section 3 we have shown that the sum of the stress-energy tensors of all the four fields inside the body is zero. With this in mind, the metric tensor components were calculated as functions of the current radius in Eq. (27). As a result, on the surface of the body at $r = a$ it became possible to compare the metric inside and outside the body and to determine the unknown coefficient A_3 in the external metric (30).

The metric tensor components g_{00} and g_{11} outside the fixed spherical body in the covariant theory of gravitation (CTG), that were presented in Eq. (30), were specified by us in Eqs. (33) and (34). It turns out that these components are the functions of the scalar potentials of all the fields, so that, for example, the pressure field inside the body also influences the metric outside the body. However, the main contribution to the metric is made by the scalar potential of the gravitational field $\psi = -\frac{Gm_g}{r}$. Apparently this is due to the fact that the expression for the scalar potential ψ includes

the gravitational mass m_g that characterizes the source of the field and the gravitation force. At the same time the relativistic energy is proportional to the inertial mass M , while for an external observer the mass M is the rest mass and characterizes the system with respect to the forces acting on it. Both of these masses differ from each other by the mass-energy of the particles' binding by means of the fields [26]. As for the electromagnetic field, its contribution is secondary. The body's charge is only indirectly included in the rest mass of the body and is not directly included in the gravitational mass. The electric field potentials vanish in neutral bodies in Eq. (34). Thus, the gravitational field is the main factor that distinguishes the curved space-time metric from the Minkowski flat one.

Our calculations allowed us to calculate the metric CTG inside the body and to refine the metric outside the body, but there was one more unknown adjustable coefficient β in the metric tensor components. Its appearance can be due to the assumption that the coefficient (29) has an exact value, so that the coefficient β is intended to ensure the correct value of the metric. The value of the coefficient β can be determined in the gravitational experiments, in which the space-time metric should be taken into account.

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THE AUTHOR

FEDOSIN Sergey G.

22–27, Architect Sviyazev St., Perm, Perm Territory, 614088, Russian Federation
sergey.fedosin@gmail.com



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СВЕДЕНИЯ ОБ АВТОРЕ

ФЕДОСИН Сергей Григорьевич – физик-исследователь, г. Пермь, Пермский край, Российская Федерация.

614088, Российская Федерация, г. Пермь, Пермский край, ул. Архитектора Свизева, 22, кв. 79.
sergey.fedosin@gmail.com